# Origin Space Telescope Mission Concept

- JWST sized telescope (~25 m<sup>2</sup>, 5.9 m), diffraction limited @ 30 μm
- Launch 2035 on large rocket (SLS or BF3)
- Spitzer-like architecture with minimal deployments
- Wavelength Coverage 3-600 µm
- Cold (~4.5 K) telescope with long-life cryo-coolers
- Three cold (~4.5 K) modular instruments for easy serviceability
- · Follows NASA standard test-as-you-fly philosophy
- Efficient mapping: up to 60" per second
- Technology development: detectors
- Mission operations at Sun-Earth L2 orbit
- 5 year lifetime, 10 year goal

10<sup>12</sup>

m<sup>-2</sup> (hrs) 10<sup>1</sup>

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to 5σ

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5





Instrument	Wavelengths (µm)	Instantaneous Field of View	Observing Modes
<b>MISC</b> Mid-Infrared Spectrometer	2.8 - 20	5" (2.8 - 10.5 μm) 3."4 (10.5 - 20 μm) set by field stop	Transiting exoplanet spectrometer
<b>OSS</b> Origins Survey Spectrometer	25 - 590	14' slit (grating mode); single beam (high-resolution modes)	R = 300 grating spectroscopy with wide-field survey capability R = 43,000(112 $\mu$ m/ $\lambda$ ) FTS R = 325,000(112 $\mu$ m/ $\lambda$ ) Etalon
<b>FIP</b> Far-Infrared	50 and 250	3.'6 x 2.'5 (50 μm) 13.'5 x 9' (250 μm)	Broadband imaging, pointed observations or wide-field survev:

Origin Space Telescope Instruments



polarimetry.





Spatial-Spectral Survey Time

SPICA

ALMA

Herschel PACS

instruments that provide low (FIP) or moderate

**SPICA** - LRS

previous and planned missions, with

spectral resolution (OSS).

Polarimeter

National Aeronautics and Space Administration



Learn more at origins.ipac.caltech.edu

## Tracing galaxy and black hole growth across cosmic time

Origins will measure how galaxies form stars, build-up metals, and grow their supermassive black holes from reionization to today

With deep and wide 3D surveys of millions of galaxies, Origins will:

• measure star formation and black hole accretion rates over 95% of cosmic time

 trace the rise of metals, dust and organic molecules

• measure galactic outflows and feedback over the past 10 Gyr

Origins will detect nearly

disks to reveal the trail of

life's ingredients through

all evolutionary stages from planetary origins to

habitable worlds.

the Galaxy.

planet.

Origins will use the ground-state line of

deuterated hydrogen

(HD) to determine the

planet-forming mass in

Origins will measure the

D/H ratio in over 100

comets to understand

our own inhabited

the delivery of water to

the entire rotational

spectrum of water in

1000 planet-forming



Searching for biosignatures in nearby exoplanets

With a mid-infrared transit spectrometer operating simultaneously over a wide wavelength range, OST will assess the habitability of nearby exoplanets and search for signs of life.

By constraining the presence of bio-indicators (H<sub>2</sub>O and CO<sub>2</sub>) and biosignatures  $(O_3, N_2O, and CH_4)$  in rocky planets transiting M dwarfs (the most common type of star in our galaxy), OST will be capable of answering the age-old questions of "Are we alone?"



# Following planetary system formation

Water in planet-forming

disks

from the interstellar medium to life-bearing worlds

Origin of water

### **Discovery space** of Origins

Origins will make significant progress on the three science themes within one year of time leaving 4+ years for all science. Origins is not only capable of addressing known questions but has a vast discovery space that will enable astronomers in the 2030s to find new phenomena and address unknown questions. All science programs on Origins will be selected by the community via peer review.

What science will you pursue with Origins?

Delivery of water

Water in planetary

systems

